Hamiltonian symmetries of the gauge theories in the light front

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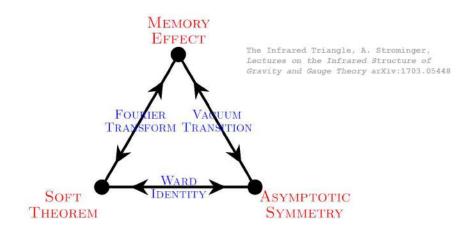
- 1 Infrared structure of gauge theories
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- 6 Discussion

IR region of theories with massless particles in asymptotically flat spaces



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Motivation

Hamiltonian treatment of asymptotic symmetries

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[Bondi, van der Burg, Metzner 1962; Sachs 1962]
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BMS symmetry – infinite-dimensional asymptotic symmetry at the null boundary of 4D asymptotically flat spacetimes

Celestial holography

Duality between massless 4D asymptotically flat spacetimes and 2D CFT on the celestial sphere

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Duality between massless 4D asymptotically flat spacetimes and 2D CFT on the celestial sphere

- Asymptotic symmetries in electromagnetism and Yang-Mills theory
- 2D realization of soft symmetries in electromagnetism

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[He, Mitra, Porfyriadis, Strominger 2014;
Nande, Pate, Strominger 2018]
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Extension to Yang-Mills theory

[Strominger 2014; He, Mitra, Strominger, 2016]

Vacuum degeneracy in gauge theories $(\omega \to 0)$

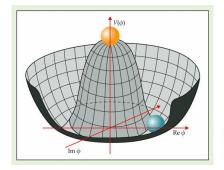
 \Leftrightarrow Enhancement of symmetries at the boundary of flat spacetime $(r \to \infty)$

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 - Vacuum state $e^{S[\eta]}\ket{A}=\ket{A+\eta}$ \Leftrightarrow $\delta\ket{A}=\eta$

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Spontaneous symmetry breaking, J. Lykken, M. Spiropulu, The future of the Higgs boson, Physics Today 66, 12, 28 (2013)

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What to expect at the null infinity?

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- Realisation of a canonical analysis in a null foliation [Dirac 1949]
- The induced metric on a null hypersurface is degenerate
- Double-null foliation = 2 + 2 formalism in GR, complicated symplectic structure, difficult to quantize [d'Inverno, Smallwood 1980]
- Ashtekar variables in GR, simpler symplectic structure, but still difficult to quantize [Ashtekar 1986, 1987]

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Peculiarities of the light-front dynamics in the Minkowski space

- **Light-cone coordinates** $x^{\pm} = \frac{x^0 \pm x^3}{\sqrt{2}}$; Time coordinate $u = x^-$
- Increased number of isometries of the surface u = const. compared to $t = x^0 = \text{const.}$ (one more because of degenerated direction)

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- Increased number of isometries of the surface u = const. compared to $t = x^0 = \text{const.}$ (one more because of degenerated direction)
- Dispersion equation for a massive scalar

$$p^2 = m^2$$
 \Rightarrow Energy $E = p^- = \frac{(p^\perp)^2 + m^2}{2p^+}$

 \Rightarrow Consequences: $p^+>0$ and trivial physical vacuum, $p_{\rm vac}^{\mu}=0$

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- Light-cone actions are first order in velocities [Steinhardt 1980]

Kinetic term
$$T=-rac{1}{2}\,\left(\partial\phi
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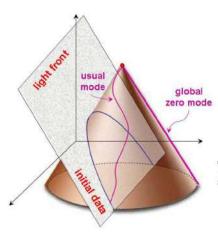
- \Rightarrow The canonical momentum $\pi = \partial_+ \phi$ is not invertible
- \star New constraint $\chi \equiv \pi \partial_+ \phi \approx 0$
- \bigstar It does not commute with itself, $\{\chi(x),\chi(x')\}_{u=u'}=-2\partial_+\delta(x-x')$
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 - Global zero mode in massless theories [Alexandrov, Speziale 2015]
- A massless particle worldline is parallel to the light front hypersurface, not determined by the initial data
- It has vanishing energy, $E=P^- o 0$ (Fourier momenta $P^-=0$, $P^\perp=0$)

Global zero mode



Global zero mode, First order gravity on the light front, S. Alexandrov, S. Speziale, Phys.Rev.D 91 (2015) 6, 064043

Null foliated reference frame

• Minkowski metric in D = 4 in the spherical coordinates (t, r, y^A)

$$M_4: ds^2 = -dt^2 + dr^2 + r^2 d\Omega^2$$

$$\mathbb{S}^2$$
: $d\Omega^2 = \gamma_{AB}(y) dy^A dy^B$

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 S^2 : $d\Omega^2 = \gamma_{AB}(y) dy^A dy^B$

• Time coordinate $u = t - \epsilon r$, $-1 \le \epsilon \le 1$

 $\epsilon = 1$ retarded time

 $\epsilon=0$ proper time of a massive particle

 $\epsilon = -1$ advanced time

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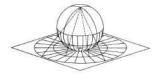
Stereographic projection, T. Apostol, Mathematical Analysis (1973)



 $\epsilon = 1$ retarded time

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• Coordinates on \mathbb{S}^2 : stereographic projection $(\theta, \varphi) \to y^A = (z, \bar{z})$

$$z=\mathrm{e}^{\mathrm{i} arphi}\cot rac{ heta}{2}$$
 , $ar{z}=\mathrm{e}^{-\mathrm{i} arphi}\cot rac{ heta}{2}$



• Minkowski metric $\mathfrak{g}_{\mu\nu}$ in the coordinates $x^{\mu}=(u,r,y^A)$:

$$ds^{2} = -du^{2} - 2\epsilon dudr + (1 - \epsilon^{2}) dr^{2} + r^{2}d\Omega^{2} \qquad \sqrt{\mathfrak{g}} = r^{2}\sqrt{\gamma}$$

ullet \mathbb{S}^2 metric in the complex coordinates

$$\gamma_{AB} = \left(egin{array}{cc} 0 & \gamma_{zar{z}} \ \gamma_{zar{z}} & 0 \end{array}
ight), \quad \gamma_{zar{z}} = rac{2}{(1+zar{z})^2} = \sqrt{\gamma}$$



 \bullet Electromagnetic action in the background $\mathfrak{g}_{\mu\nu}$

$$I[A] = -rac{1}{4e^2}\int \mathrm{d}^4x\,\sqrt{\mathfrak{g}}\,F^{\mu\nu}F_{\mu\nu} \qquad (F_{\mu\nu} = \partial_\mu A_
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In components:

$$\begin{array}{lll} \pi^{u} &= 0 & \dot{A}_{u} & \times \\ \pi^{r} &= \frac{r^{2}}{e^{2}} \sqrt{\gamma} \, F_{ur} & \dot{A}_{r} & \sqrt{} \\ \pi^{A} &= -\frac{1}{e^{2}} \sqrt{\gamma} \gamma^{AB} \left[(\epsilon^{2} - 1) F_{uB} - \epsilon F_{rB} \right] & \dot{A}_{B} & ? \end{array}$$

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- The limit $\epsilon^2 \to 1$ is discontinuous
- The action in the light-cone ($\epsilon^2 = 1$) has an additional constraint

In the Bondi reference frame ($\epsilon^2 = 1$)

• Primary constraints

$$\pi^u \approx 0$$
, $\chi^A \equiv \epsilon \pi^A - \frac{1}{e^2} \sqrt{\gamma} \gamma^{AB} F_{rB} \approx 0$

In the Bondi reference frame ($\epsilon^2 = 1$)

- Primary constraints $\pi^u \approx 0$, $\chi^A \equiv \epsilon \pi^A \frac{1}{\epsilon^2} \sqrt{\gamma} \gamma^{AB} F_{rB} \approx 0$
- Total Hamiltonian [Dirac 1964]

$$\mathcal{H}_{T} = \frac{e^{2}(\pi^{r})^{2}}{2r^{2}\sqrt{\gamma}} + \frac{e^{2}\tilde{\pi}_{A}\pi^{A}}{2\sqrt{\gamma}} + \frac{\sqrt{\gamma}}{4e^{2}r^{2}}\tilde{F}^{AB}F_{AB} - A_{u}\partial_{i}\pi^{i} + \lambda_{u}\pi^{u} + \lambda_{A}\chi^{A}$$

- Hamiltonian multipliers $\lambda_u(x)$, $\lambda_A(x)$ incorporate constraints

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- Hamiltonian multipliers $\lambda_u(x)$, $\lambda_A(x)$ incorporate constraints
- Canonical Poisson brackets $\left\{A_{\mu}(x),\pi^{\nu}(x')\right\}_{u=u'}=\delta^{\nu}_{\mu}\,\delta^{(3)}(x-x')$
- Evolution $\dot{\Phi}(x) = \{\Phi(x), H_T\}$

Symplectic matrix

$$\left\{ \chi^{A}(x), \chi^{B}(x') \right\} = \Omega^{AB}(x, x') \equiv -\frac{2\epsilon}{e^{2}} \sqrt{\gamma} \gamma^{AB} \partial_{r} \delta^{(3)}$$

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- If Ω^{AB} is invertible: χ^A are second class (eliminate redundant fields)

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- Ω^{AB} is invertible because γ^{AB} is invertible $\Rightarrow \chi^{A}$ are second class
- ullet Reduced phase space $\chi^A=0$ [Goldberg 1991, Majumdar 2022]

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Second possibility

- Ω^{AB} is invertible, but its inverse is not unique
- Ω^{AB} is infinite-dimensional matrix and it has zero modes

$$\int d^3x'\,\Omega^{AB}\,V_B' = -\tfrac{2\varepsilon}{\varepsilon^2}\,\sqrt{\gamma}\gamma^{AB}\partial_r V_B = 0 \quad \Rightarrow \quad \underset{\scriptscriptstyle \bullet \, \square \, P}{\longrightarrow} \quad \underset{$$

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Symplectic matrix

$$\left[\left\{ \chi^{A}(x), \chi^{B}(x') \right\} = \Omega^{AB}(x, x') \equiv -\frac{2\epsilon}{e^{2}} \sqrt{\gamma} \gamma^{AB} \partial_{r} \delta^{(3)} \right]$$

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$$\Rightarrow \chi^{A}_{(0)}(y)$$
 is first class constraint (r-independent part of the constraint)

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Consistency conditions

Consistency conditions

Conservation of constraints during their evolution

$$\dot{\pi}^u = 0 \quad \Rightarrow \quad \chi = \partial_i \pi^i \approx 0 \quad \text{(differential Gauss law)}$$

$$\dot{\chi}^A = 0$$
 \Rightarrow differential equation in the multiplier

Consistency conditions

Conservation of constraints during their evolution

$$\dot{\pi}^u = 0$$
 $\Rightarrow \chi = \partial_i \pi^i \approx 0$ (differential Gauss law) $\dot{\chi}^A = 0$ \Rightarrow differential equation in the multiplier

• The multiplier λ_A is not fully determined

$$\partial_r \lambda_A = -\frac{\epsilon e^2}{2\sqrt{\gamma}} \partial_r \tilde{\pi}_A - \frac{1}{2r^2} \nabla^B F_{AB} + \frac{\epsilon e^2}{2r^2} \partial_B \left(\frac{\pi^r}{\sqrt{\gamma}} \right)$$
 $\lambda_A = \bar{\lambda}_A + \Lambda_A(y)$

- A free function $\Lambda_A(y)$ is due to the zero modes of Ω^{AB}
- $\bar{\lambda}_A$ determined part of λ_A



Summary of the constraints

Primary constraints:
$$\pi^u$$
, $\chi^A = \epsilon \pi^A - \frac{1}{e^2} \sqrt{\gamma} \gamma^{AB} F_{rB}$

Secondary constraint: $\chi = \partial_i \pi^i$.

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Nature of the constraints

Summary of the constraints

Primary constraints:
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, $\chi^A = \epsilon \pi^A - \frac{1}{e^2} \sqrt{\gamma} \gamma^{AB} F_{rB}$

Secondary constraint: $\chi = \partial_i \pi^i$.

Nature of the constraints

- π^u first class, A_u is a multiplier in the Hamiltonian
- χ first class, differential Gauss law, $\pi^i = \sqrt{\gamma} \, {\sf E}^i$
- $\chi_{(0)}^A$ first class, r-independent part of the constraint
- $\chi^A_{(n)}$ $(n \ge 1)$ second class, coefficients of the Taylor expansion in 1/r
- We have to expand all the fields asymptotically in the vicinity of the boundary r = const → ∞.

Standard asymptotic conditions of the fields [Strominger 2014]

$$A_{u} = \mathcal{O}(\frac{1}{r}), \qquad A_{r} = \mathcal{O}(\frac{1}{r^{2}}), \qquad A_{A} = \mathcal{O}(r^{0}),$$

$$\pi^{u} = 0, \qquad \pi^{r} = \mathcal{O}(r^{0}), \qquad \pi^{A} = \mathcal{O}(\frac{1}{r^{2}})$$

• Boundary fields: $A_{(0)A}$, $\pi_{(0)}^r$

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• Boundary fields: $A_{(0)A}$, $\pi^r_{(0)}$

Summary

1^{st} class constraints	Parameters	Generators	Charges
π^u , $\chi=\partial_i\pi^i$ $\chi^A_{(0)}$	$\theta_u, \ \theta \\ \eta_A(y)$	$G[heta] \ S[\eta]$	$egin{aligned} Q[heta] \ Q_s[\eta] \end{aligned}$

• Smeared generators

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• $G[\theta]$ generates standard gauge transformations, $\delta A_{\mu} = -\partial_{\mu}\theta$, because it is satisfied $\theta_{\mu} = -\dot{\theta}$ [Castellani 1974]

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- New symmetry generator $S[\eta]$ has an r-independent parameter $\eta_A(y)$ (to eliminate $\chi^A_{(n)}$, $n \ge 1$)
- Transformation law of the fields

$$egin{array}{lll} \delta_{ heta}A_{\mu} &= -\partial_{\mu} heta\,, & \delta_{\eta}A_{\mu} &= \epsilon\,\eta_{A}\,\delta_{\mu}^{A} \ & & \delta_{ heta}\pi^{\mu} &= 0\,, & \delta_{\eta}\,\pi^{\mu} &= rac{1}{e^{2}}\,\delta_{r}^{\mu}\sqrt{\gamma}\,
abla_{A}\eta^{A} \end{array}$$

- Improper transformations: $\theta_{(0)}$, η^A [Benguria, Cordero, Teitelboim 1977]
- \bigstar η^A acts on the boundary fields only

$$\delta_{\eta}A_{(0)A}=\epsilon\,\eta_A$$
 , $\delta_{\eta}A_{(n)A}=0$, $n\geq 1$ (similarly for π^r)

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Improved generators and charges

Improved generators

$$egin{array}{ll} G_Q[heta] &= G[heta] + Q[heta] & (Q[heta] &= {
m surface\ term}) \ & S_Q[\eta] &= S[\eta] + Q_s[\eta] & (Q_s[heta] &= {
m surface\ term}) \end{array}$$

Differentiability

$$\begin{split} \delta G_Q[\theta] &= \int d^3x \, \left(\frac{\delta G_Q[\theta]}{\delta A_\mu} \, \delta A_\mu + \frac{\delta G_Q[\theta]}{\delta \pi^\mu} \, \delta \pi^\mu \right) \\ \delta S_Q[\eta] &= \int d^3x \, \left(\frac{\delta S_Q[\eta]}{\delta A_\mu} \, \delta A_\mu + \frac{\delta S_Q[\eta]}{\delta \pi^\mu} \, \delta \pi^\mu \right) \end{split}$$

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Charges

$$Q[\theta] = -\oint d^2y \,\theta \,\pi^r$$

$$Q_s[\eta] = \frac{1}{e^2} \oint d^2y \,\sqrt{\gamma} \,\eta^A A_A$$

- Infinite number of global charges (Laurent coefficients).

Charge algebra

- Reduced phase space: $G_Q[\theta] = Q[\theta], S_Q[\eta] = Q_s[\eta]$
- Abelian charge algebra

$$\{Q[\theta_1], Q[\theta_2]\} = 0$$

 $\{Q_s[\eta_1], Q_s[\eta_2]\} = 0$
 $\{Q[\theta], Q_s[\eta]\} = C[\theta, \eta]$

• Central charge $C[\theta,\eta]=rac{1}{e^2}\oint \mathrm{d}^2y\,\sqrt{\gamma}\,\eta^A\partial_A\theta
eq 0$

Charge algebra

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- ullet Holographic conjugate pairs on \mathbb{S}^2 [Donnay, Puhm, Strominger 2019]

$$\{Q[\theta], Q_s[\eta]\} = C[\theta, \eta] \quad \leftrightarrow \quad \{q, p\} = 1$$

 $Q[\theta]$ – conformally soft photon mode

$$Q_s[\eta]$$
 – Goldstone current



Mode expansion of the charge algebra

Laurent series

$$\psi(z,\bar{z}) = \sum_{n=-\infty}^{+\infty} \sum_{m=-\infty}^{+\infty} \frac{\psi_{nm}}{z^{n+h}\bar{z}^{m+\bar{h}}}$$

- The powers (h, \bar{h}) are related to the spin of the tensor ψ
- Scalars $\pi^r:(0,0)$
- Vectors $A_{z}:(1,0), \quad A_{\bar{z}}:(0,1)$

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- Scalars $\pi^r:(0,0)$
- Vectors $A_z:(1,0), A_{\bar{z}}:(0,1)$
- Charges

$$Q[\theta] = \sum_{n,m} \theta_{nm} G_{nm} \in \mathbb{R}$$

$$Q_s[\eta] = \sum_{n,m} (\eta_{nm} \bar{S}_{nm} + \bar{\eta}_{nm} S_{nm}) \in \mathbb{R}$$

Generators

$$G_{nm} = 4\pi^2 \, \pi_{1-n,1-m}$$
 $S_{nm} = -\frac{4\pi^2}{e^2} \, A_{-n,-m}$
 $\bar{S}_{nm} = -\frac{4\pi^2}{e^2} \, \bar{A}_{-n,-m}$

Algebra (non vanishing brackets only)

```
\{G_{nm}, S_{kl}\} = \kappa n \delta_{n+k,0} \delta_{m+l,0}
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• Level of the algebra: $\kappa = \frac{4\pi^2}{e^2}$

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- Level of the algebra: $\kappa = \frac{4\pi^2}{e^2}$
- Change of the basis: $(G_{nm}, S_{nm}, \bar{S}_{nm}) \rightarrow (R_{nm}, J_{nm}, \bar{J}_{nm})$
- Generalization of the Kac-Moody algebra

Abelian Kac-Moody subalgebras

• We obtain six Abelian KM algebras $\{j_n,j_m\}=\kappa n\,\delta_{n+m,0}$

Currents
$$j_n$$
 Levels
$$J_{n0}, J_{0n} \qquad \kappa, -\kappa$$

$$\bar{J}_{n0}, \bar{J}_{0n} \qquad -\kappa, \kappa$$

$$R_{n0}, R_{0n} \qquad \kappa, \kappa$$

- Non vanishing mixed brackets: $\{R_{n0},J_{m0}\}$, $\{R_{0n},\bar{J}_{0m}\} \neq 0$
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- Each KM algebra is naturally generated by a current that is a holomorphic or anti-holomorphic function.
- $\{J_{00}, \bar{J}_{00}, R_{00}\}$ span the global Abelian algebra (only two independent in EM)

Beyond U(1) – conformal symmetry

- Conformal plane a realization of conformal symmetry described by Virasoro algebra
- Virasoro algebra obtained from KM algebra using the Sugawara construction [Sugawara 1967]

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- In progress: Relation to the global 4D Poincaré generators.

Yang-Mills theory

$$I[A] = -rac{1}{4g^2}\int \mathrm{d}^4x \sqrt{\mathfrak{g}}\,F_a^{\mu\nu}F_{\mu\nu}^a$$

Constraints

$$\pi_a^u \approx 0 \,, \quad \chi_a^A \equiv \epsilon \, \pi_a^A - \tfrac{1}{g^2} \, \sqrt{\gamma} \gamma^{AB} \, F_{rB}^a \approx 0 , \quad \chi_a \equiv D_i \pi_a^i \approx 0 \,$$

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Constraint algebra

$$\begin{cases} \chi_{a}, \chi_{b}' \rbrace &= f_{ab}^{\ c} \chi_{c} \delta^{(3)} \\ \chi_{a}, \chi_{b}'^{A} \rbrace &= f_{ab}^{\ c} \chi_{c}^{A} \delta^{(3)} \\ \chi_{a}^{A}, \chi_{b}'^{B} \rbrace &= \Omega_{ab}^{AB}(x, x')$$

• Non-Abelian symplectic matrix

$$\Omega_{ab}^{AB}(x,x') = -\frac{2\epsilon}{g^2} \sqrt{\gamma} \gamma^{AB} \left(g_{ab} \partial_r + f_{abc} A_r^c \right) \delta^{(3)}$$



Zero mode

$$\partial_r V_A = -[A_r, V_A] \quad \Rightarrow \quad V_A(x) = U^{-1} V_{(0)A}(y) U,$$
 with $U = \exp\left(-\int_r^\infty \mathrm{d} r \, A_r\right)$ and the bdy. condition $V_A|_{r \to \infty} = V_{(0)A}(y)$

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Charges

$$Q[\theta] = -\oint \mathrm{d}^2 y \, heta^a \pi_a^r$$
, $Q_s[\eta] = rac{1}{g^2} \oint \mathrm{d}^2 y \, \sqrt{\gamma} \, \eta_A^a A_a^A$

Local transformations

$$\begin{split} \delta_{\theta,\eta}A_u^a &= \theta_u^a & \delta_{\theta,\eta}A_r^a = -D_r\theta^a \\ \delta_{\theta,\eta}A_A^a &= -D_A\theta^a + \epsilon\,\eta_A^a \end{split}$$

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Non-Abelian charge algebra

Mode algebra

$$\left\{ G_{nm}^{a}, G_{kl}^{b} \right\} = f_{c}^{ab} G_{n+k,m+l}^{c}$$

$$\left\{ G_{nm}^{a}, S_{kl}^{b} \right\} = f_{c}^{ab} S_{n+k,m+l}^{c} + \kappa n g^{ab} \delta_{n+k,0} \delta_{m+l,0}$$

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- Level $\kappa = \frac{4\pi^2}{g^2}$
- One can apply the Sugawara method again...
- Symmetries at the asymptotic null boundary, described by KM algebras and Virasoro algebras, are general features of 4D gauge theories

Degrees of freedom count

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• Dirac formula d.o.f. = $N - N_{1^{st}class} - \frac{1}{2} N_{2^{nd}class}$

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- d.o.f. = $4 2 \frac{1}{2}2 = 1$ **WRONG!!** d.o.f. = 2
- The Dirac formula is applicable only when the multipliers are either arbitrary (1st class constraints) or fully determined (2nd class constraints).
- It fails when the multipliers satisfy a differential equation.
- In the nul foliation: $\partial_r \lambda^A = f^A \quad \Rightarrow \quad \lambda^A = \Lambda^A(y) + \bar{\lambda}^A$



Asymptotic conditions

- Invariance of boundary conditions under Poincaré transformations is not straighforward
- Hamiltonian treatment at spatial infinity needs additional parity conditions to ensure invariance under boosts
- Electromagnetism [Henneaux, Troessaert 2018]
- Yang-Mills [Tanzi, Giulini 2020]
- Null-slices foliated standard b.c. in electromagnetism are invariant under Poincaré group. [Bunster, Gomberoff, Pérez 2018]
- We showed the Poincaré invariance in the non-Abelian case.

Poincaré transfromations

- We found several Kac-Moody algebras, but not all of them are related to the global Poincaré symmetry in 4D spacetime.
- We constructed a generator of 4D Poincaré transformations and its action at the light front, by writing the YM stress tensor in the canonical form,

$$T^{\mu}_{\ \nu} = rac{1}{g^2}\,\left(F^{\mu\alpha}_aF^a_{\nu\alpha} - rac{1}{4}\,\delta^{\mu}_{
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We are working on showing its relation with the Virasoro generators.

To be done

- Hamiltonian treatment of the gravitational action using the null foliation.
- · Description of a holographic theory.
- Addition of the θ -term in the action (Pontryagin topological invariant with the couplig θ), which will change the central charges.

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THANK YOU!

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Black holes and asymptotic symmetries

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